

Processing HDF to FITS Image: Python Pipeline Mode

¹Alexander Akoto-Danso, ²Felix Tetteh Madjitey, ³Emmanuel Proven-Adziri, ⁴Theophilus Ansah-Narh, ⁵Marcellin Atemkeng

^{1,2,3}Ghana Space Science and Technology Institute (GSSTI), Ghana Atomic Energy Commission (GAEC).

^{4,5}RATT, Rhodes University, South Africa.

Abstract - Research scientists and students especially those in astronomy normally use image data in Flexible Image Transport System (FITS) format for their simulations. Simulation tools like Meqtrees, Oskar, Map simulator and Swarp all run using FITS file. Sometimes, the image data is stored in different formats such as Hierarchical Data Format (HDF) and other formats. We need to convert the data into FITS form before using for simulations. The paper therefore presents a python program on how to extract data from HDF into FITS file. The h5py helps us to extract the image data from HDF, numpy then converts the data into an array and pyfits helps us to write the extracted data in FITS format. The h5py, numpy and pyfits are imported into python for the conversion to take place. The developed code and output are presented in Appendices A and B respectively.

Keywords - Hierarchical Data Format (HDF), Flexible Image Transport System (FITS), Python, h5py, numpy, pyfits.

1. Introduction

Hierarchical Data Format (HDF) was formally developed at the National Center for Supercomputing Applications with its former version being HDF4 and now changed to HDF5. It is designed to store any large amount of scientific data in an organized form. The current version, HDF5 simplifies the file structure into two major types of objects:

- *Datasets*:- these are multidimensional arrays of a homogenous type.
- *Groups*:- these are structures in a form of containers that hold the datasets and other groups.

Detailed explanations on how to use HDF format are [1 and 2].

Most simulation tools such as MeqTrees [3], OSKAR [4], MIT Array Performance Simulator (MAPS) [5] and SWarp [6] all run using FITS file. Sometimes, the image

data is stored in different formats such as HDF and other formats and we need to convert the data into FITS form before using for simulations.

FITS is also used to store both image and non-image data, such as spectra, photon lists, data cubes and multi-table databases. It is one of the commonest file format used in astronomy. Detailed explanations on how to use FITS file are [7, 8 and 9].

2. Program Structure

Figure 1 is a flow chart describing clearly how HDF is converted into FITS file. The h5py, numpy and pyfits are python modules or packages that must be installed to be able to import them into the python platform for conversion.

Appendix A shows the python script developed to convert HDF to FITS file.

3. Results and Discussions

Appendices B – 1 to B – 4 show the outputs of the FITS file created. Appendix B – 1 shows the running process during the conversion by going through the flowchart in Fig. 1. Appendix B – 2 shows the output of the headers when no update is made and appendix B – 3 is the updated output of the headers using World Coordinate System (WCS) [10] which can always be updated to suit the required image data.

Appendix B – 4 is the image data in array form extracted into FITS file. The FITS file created consists of more *header* and *data units*, normally referred to as HDUs. The HDU consists of the header which describes the data contained in it, followed by the data itself.

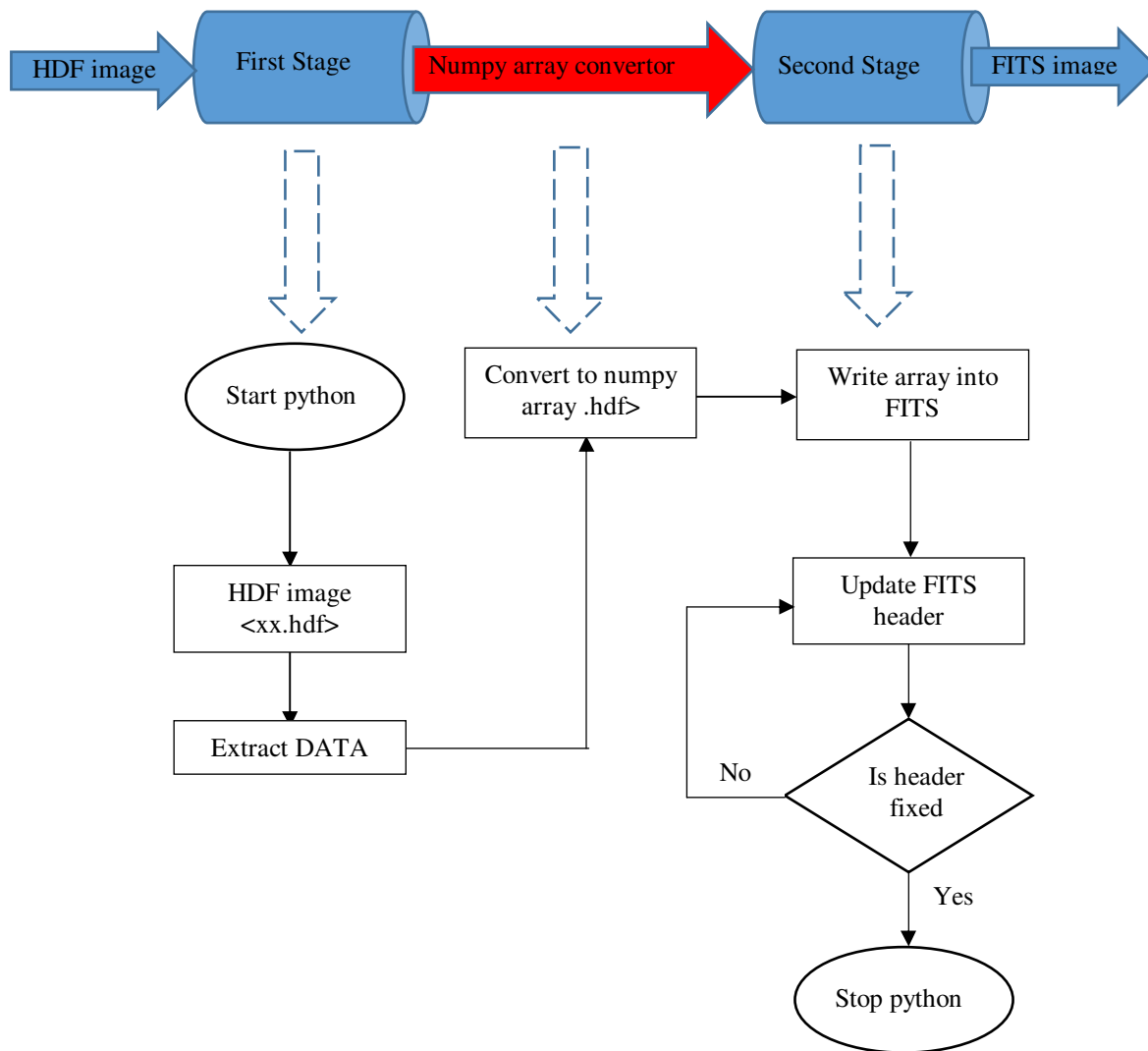


Fig. 1: Procedure for Converting HDF to FITS

4. Conclusion and Future Work

In this paper, we provided how to convert HDF to FITS using python module pyfits. Clearly, the paper showed the basic operations with FITS file such as *reading, modifying and writing an image data*.

The FITS file created in this paper is the Primary HDU and therefore, a continuation of this study is to develop a *Multi – Extension FITS (MEF) file in BINTABLE format*. This would be very important particularly in the application of *Hierarchical, Equal Area and Iso-Latitude Pixelisation (HEALPIX)* for spherical designs such as sky maps.

The National Center for Supercomputing Applications (NCSA) HDF group and others are currently working together to convert all HDF objects to FITS and vice versa, making it possible to access FITS files with HDF software and vice versa [11] but until then this paper should easily help in converting HDF to FITS.

Appendix A

```

#!/usr/bin/env python
"""
hdf2fits.py
=====
Converts HDF to FITS format
"""
__version__ = "1.0"
import h5py
    
```

```
import pyfits
import numpy as np
from pyfits import getdata, getheader
print "-----"
print "\n >> Please enter HDF5 file name to convert
(including extension):"
print "-----"
filename=raw_input() #Input .hdf5 file
print ">> Opening file and reading data"
print "-----"
hdf_data =h5py.File(filename,'r')
print ">> Extracting data from HDF file"
print "-----"
extract_data = hdf_data.values()[:]
print ">> Converting HDF data into arrays"
print "-----"
a_ray = np.array(extract_data)
#####
#####
print ">> Converting HDF data into FITS"
print "-----"
hdu = pyfits.PrimaryHDU(a_ray)
hdu.writeto('newhd2fits.fits')
#####
#####
print "-----"
print "\n >> Please checking & fixing headers:"
print "-----"
filename1='newhd2fits.fits' #Input .fits file

print ">> Opening FITS file and reading header"
print "-----"

hdulist = pyfits.open(filename1)
in_header=hdulist[0].header
in_header.update('BITPIX', -64)
print ">> BITPIX Fixed\n >> Now checking for all
required fits header keywords"
print "-----"

try:
    in_header['COMMENT']
except KeyError:
    print "\t No COMMENT Found, adding entry for this"
    in_header.update('COMMENT', 'Standard WCS
reduction:')
try:
    in_header['BZERO']
except KeyError:
    print "\t No BZERO Found, adding entry for this"
    in_header.update('BZERO',0.0,'PhysValue = BZERO +
BSCALE * ArrayValue')
try:
```

```
    in_header['BSCALE']
except KeyError:
    print "\t No BSCALE Found, adding entry for this"
    in_header.update('BSCALE',1.0,'PhysValue = BZERO
+ BSCALE * ArrayValue')

try:
    in_header['CRVAL1']
except KeyError:
    print "\t No CRVAL1 Found, adding entry for this"
    in_header.update('CRVAL1',11.88798983,'WCS Ref
value (RA in decimal degrees)')
try:
    in_header['CRVAL2']
except KeyError:
    print "\t No 'EPOCH' Found, adding entry for this"
    in_header.update('EPOCH',1999.55329832624,'Epoch
in Julain Year at start of observation')
try:
    in_header['CRPIX1']
except KeyError:
    print "\t No 'CRPIX1' Found, adding entry for this"
    in_header.update('CRPIX1',256.5,'WCS Coordinate
reference pixel')
try:
    in_header['CRPIX2']
except KeyError:
    print "\t No 'CRPIX2' Found, adding entry for this"
    in_header.update('CRPIX2',356.5,'WCS Coordinate
reference pixel')
try:
    in_header['CD1_1']
except KeyError:
    print "\t No 'CD1_1' Found, adding entry for this"
    in_header.update('CD1_1',-
0.000277777783923599,'WCS Coordinate scale matrix')
try:
    in_header['CD1_2']
except KeyError:
    print "\t No 'CD1_2' Found, adding entry for this"
    in_header.update('CD1_2',1.78947724260645E-08,
'WCS Coordinate scale matrix')
try:
    in_header['CD2_1']
except KeyError:
    print "\t No CD2_1 Found, adding fake entry for this"
    in_header.update('CD2_1',1.78947724260645E-
08,'WCS Coordinate scale matrix')
try:
    in_header['CD2_2']
except KeyError:
    print "\t No CD2_2 Found, adding fake entry for this"
```

```
in_header.update('CD2_2',0.000277777783923599,'WCS
Coordinate scale matrix')
try:
    in_header['CD3_1']
except KeyError:
    print "\t No 'CD3_1' Found, adding entry for this"
    in_header.update('CD3_1',1.78947724260645E-08,
'WCS Coordinate scale matrix')
try:
    in_header['CD3_2']
except KeyError:
    print "\t No 'CD3_2' Found, adding entry for this"
    in_header.update('CD3_2',1.78947724260645E-08,
'WCS Coordinate scale matrix')
try:
    in_header['CTYPE1']
except KeyError:
    print "\t No CTYPE1 Found, adding entry for this"
    in_header.update('CTYPE1','RA---SIN',          'WCS
Coordinate type')
try:
    in_header['CTYPE2']
except KeyError:
    print "\t No CTYPE2 Found, adding entry for this"
    in_header.update('CTYPE2','DEC---SIN',          'WCS
Coordinate type')
try:
    in_header['EQUINOX']
except KeyError:
    print "\t No EQUINOX Found, adding entry for this"
    in_header.update('EQUINOX',2000.0,' Equinox')
try:
    in_header['CDELT1']
except KeyError:
    print "\t No CDELT1 Found, adding entry for this"
    in_header.update('CDELT1',0.0002777777845,'WCS
Coordinate scale matrix')
try:
    in_header['CDELT2']
except KeyError:
    print "\t No CDELT2 Found, adding entry for this"
    in_header.update('CDELT2',0.0002777777845,'WCS
Coordinate scale matrix')
try:
    in_header['CTYPE3']
except KeyError:
    print "\t No CTYPE3 Found, adding fake entry for this"
    in_header.update('CTYPE3','STOKES', 'axes 3 is the
spectra')
try:
    in_header['CNPIX1']
except KeyError:
```

```
    print "\t No CNPIX1 Found, adding fake entry for this"
    in_header.update('CNPIX1','0','New CNPIX1')
try:
    in_header['MJD-OBS']
except KeyError:
    print "\t No MJD-OBS Found, adding fake entry for
this"
    in_header.update('MJD-
OBS',51381.3422136574,'modified Julian date at start of
observation')
try:
    in_header['CRPIX3']
except KeyError:
    print "\t No CRPIX3 Found, adding fake entry for this"
    in_header.update('CRPIX3',1.0)
#####
try:
    in_header['CUNIT1']
except KeyError:
    print "\t No CUNIT1 Found, adding entry for this"
    in_header.update('CUNIT1','DEG', 'RA coordinate')
try:
    in_header['CUNIT2']
except KeyError:
    print "\t No CUNIT2 Found, adding entry for this"
try:
    in_header['CDELT3']
except KeyError:
    print "\t No CDELT3 Found, adding entry for this"
    in_header.update('CDELT3', 1.0)
try:
    in_header['CRPIX1']
except KeyError:
    print "\t No CRPIX1 Found, adding entry for this"
    in_header.update('CRPIX1', 256.5)
try:
    in_header['CRPIX2']
except KeyError:
    print "\t No CRPIX2 Found, adding entry for this"
#####
#####
print " >> All missing required fits header keywords
added\n >> except NAXISn. I'll check in a second."
print " -----"

data=getdata(filename1) #Read in fits data
fl_data=np.float64(data) #convert in array to float64
hdu_data=pyfits.PrimaryHDU(fl_data) #create FITS type
data

print " >> Data array Fixed"
```

```
print " -----"

naxisvals=fl_data.shape
print naxisvals[0]
print naxisvals[1]

try:
    in_header['NAXIS1']
except KeyError:
    print "\t No NAXIS1 Found, adding entry for this"
    in_header.update('NAXIS1',naxisvals[0])

try:
    in_header['NAXIS2']
except KeyError:
    print "\t No NAXIS2 Found, adding entry for this"
    in_header.update('NAXIS2',naxisvals[0])

try:
    in_header['NAXIS3']
except KeyError:
    print "\t No NAXIS3 Found, adding entry for this"
    in_header.update('NAXIS3',naxisvals[0])
print " >> NAXIS values now checked."
print " -----"
hdulist.close()
pyfits.writeto('fixed_'+filename1,fl_data,in_header)
#Write new fits file

print " >> Done!\n\n >> The corrected fits file is named
fixed_"+filename1
print " -----"
```

Appendix B – 1

```
narh@elwood:~/theo/comp_sc$ python hdf2fits.py
```

```
-----
>> Please enter HDF5 file name to convert (including
extension):
-----
pol_map.hdf5
>> Opening file and reading data
-----
>> Extracting data from HDF file
-----
>> Converting HDF data into arrays
-----
>> Converting HDF data into FITS
-----
-----
>> Please checking & fixing headers:
```

```
-----
>> Opening FITS file and reading header
-----
>> BITPIX Fixed
>> Now checking for all required fits header keywords
-----
No COMMENT Found, adding entry for this
No BZERO Found, adding entry for this
No BSCALE Found, adding entry for this
No CRVAL1 Found, adding entry for this
No 'CRVAL2' Found, adding entry for this
No 'EPOCH' Found, adding entry for this
No 'CRPIX1' Found, adding entry for this
No 'CRPIX2' Found, adding entry for this
No 'CD1_1' Found, adding entry for this
No 'CD1_2' Found, adding entry for this
No CD2_1 Found, adding entry for this
No CD2_2 Found, adding entry for this
No 'CD3_1' Found, adding entry for this
No 'CD3_2' Found, adding entry for this
No CTYPER Found, adding entry for this
No CTYPER2 Found, adding entry for this
No EQUINOX Found, adding entry for this
No CDELT1 Found, adding entry for this
No CDELT2 Found, adding entry for this
No CTYPER3 Found, adding entry for this
No CNPIX1 Found, adding entry for this
No CNPIX2 Found, adding entry for this
No RADESYS Found, adding entry for this
No MJD-OBS Found, adding entry for this
No CRPIX3 Found, adding entry for this
No CUNIT1 Found, adding entry for this
No CUNIT2 Found, adding entry for this
No CDELT3 Found, adding entry for this
>> All missing required fits header keywords added
>> except NAXISn. I'll check in a second.
```

```
-----
>> Data array Fixed
```

```
-----
1
560
>> NAXIS values now checked.
-----
>> Done!
```

```
>> The corrected fits file is named fixed_newhd2fits.fits
-----
```

Appendix B – 2

```
In [1]: import pyfits
```

```
In [2]: hdu = pyfits.open('newhd2fits.fits')
```

```
In [2]: hdu = pyfits.open('newhd2fits.fits')

In [3]: print hdu[0].header
SIMPLE =          T / conforms to FITS standard
BITPIX =         -64 / array data type
NAXIS  =          4 / number of array dimensions
NAXIS1 =        786432
NAXIS2 =          4
NAXIS3 =         560
NAXIS4 =          1
EXTEND =          T
```

Appendix B – 3

```
In [4]: hduf = pyfits.open('fixed_newhd2fits.fits')

In [5]: print hduf[0].header
SIMPLE =          T / conforms to FITS standard
BITPIX =         -64 / array data type
NAXIS  =          4 / number of array dimensions
NAXIS1 =        786432
NAXIS2 =          4
NAXIS3 =         560
NAXIS4 =          1
CRVAL1 =    11.88798983 / WCS Ref value (RA in
decimal degrees)
CRVAL2 =   -25.33718377 / WCS Ref value (DEC
in decimal degrees)
EPOCH  =  1999.55329832624 / Epoch in Julain Year
at start of observation
CRPIX1 =   256.5 / WCS Coordinate reference pixel
```

```
CRPIX2 =   356.5 / WCS Coordinate reference pixel
CD1_1  = -0.00027777778392359 / WCS Coordinate
scale matrix
CD1_2  =  1.78947724260645E-08 / WCS Coordinate
scale matrix
CD2_1  =  1.78947724260645E-08 / WCS Coordinate
scale matrix
CD2_2  =  0.000277777783923599 / WCS Coordinate
scale matrix
CD3_1  =  1.78947724260645E-08 / WCS Coordinate
scale matrix
CD3_2  =  1.78947724260645E-08 / WCS Coordinate
scale matrix
CTYPE1 = 'RA---SIN'      / WCS Coordinate type
CTYPE2 = 'DEC---SIN'    / WCS Coordinate type
EQUINOX =  2000.0 / Equinox
CDELT1 =    0.0002777777845 / WCS Coordinate
scale matrix
CDELT2 =    0.0002777777845 / WCS Coordinate
scale matrix
CTYPE3 = 'STOKES '      / axes 3 is the spectra
CNPIX1 = '0 '          / New CNPIX1
CNPIX2 = '0 '          / New CNPIX2
RADESYS = 'FK5 '       / Coordinate system
MJD-OBS =  51381.3422136574 / modified Julian date
at start of observation
CRPIX3 =    1.0
CUNIT1 = 'DEG '        / RA coordinate
CDELT3 =    1.0
COMMENT Standard WCS reduction:
```

Appendix B – 4

In [6]: print hduf[0].data

```
[[[ 1.61262884e+04 1.59303920e+04 1.55145821e+04
...,
1.18152153e+04 1.22493922e+04 1.24867062e+04]
[ 1.14865993e+01 1.89759293e+01 1.57714412e+01
...,
-5.58747410e+01 -5.59940744e+01 -5.55064362e+01]
[ 7.43384419e+01 6.80194419e+01 6.34424053e+01
...,
-3.51272863e+01 -3.31921756e+01 -3.74518863e+01]
[ 0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00
0.00000000e+00]]
[[ 1.39059342e+04 1.37400705e+04
1.33806071e+04 ...,
1.02468788e+04 1.06163441e+04 1.08212835e+04]
[ 1.23021996e+01 1.98147810e+01 1.66081076e+01
...,
-5.71332849e+01 -5.71451130e+01 -5.66618905e+01]
[7.53536529e+01 6.86724097e+01 6.38527579e+01
...,
-3.53120905e+01 -3.32438123e+01 -3.74934783e+01]
[ 0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00
0.00000000e+00]]
[[1.20725163e+04 1.19310725e+04 1.16182750e+04
...,
8.94446192e+03 9.26105631e+03 9.43923397e+03]
[1.13916453e+01 1.88206657e+01 1.57066625e+01 ...,
```

```
-5.54365238e+01 -5.53249321e+01 -5.48905659e+01]
[7.71029006e+01 7.06927265e+01 6.57704095e+01
...,
-3.64073943e+01 -3.44865497e+01 -3.87505005e+01]
[ 0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00 0.00000000e+00]]
[[ 4.93552468e-01 4.96426918e-01 4.84608608e-01
...,
4.00771575e-01 3.95754920e-01 4.08489146e-01]
[6.31610901e-02 6.13774389e-02 5.80929026e-02 ...,
-2.02320153e-02 -1.94703505e-02 -1.80827997e-02]
[3.55320610e-02 3.19743144e-02 2.87056667e-02
...,
1.36039364e-03 3.28667625e-03 2.06490153e-04]
[0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00 0.00000000e+00]]
[[4.90875403e-01 4.93740518e-01 4.81988073e-01
...,
3.98474596e-01 3.93474828e-01 4.06148925e-01]
[6.29019447e-02 6.11264250e-02 5.78342543e-02
...,
-2.03858111e-02 -1.96372520e-02 -1.82412362e-02]
[3.56680031e-02 3.21116345e-02 2.88477180e-02
...,
1.04425912e-03 2.95573873e-03 -1.24522009e-04]
[0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00 0.00000000e+00]]
[[4.88217392e-01 4.91073208e-01 4.79386150e-01
...,
3.96194360e-01 3.91211416e-01 4.03825772e-01]
[6.26412185e-02 6.08738353e-02 5.75741862e-02
...,
```

-2.05357473e-02 -1.98003038e-02 -1.83959080e-02]
[3.57999089e-02 3.22452369e-02 2.89860593e-02
...,
7.33478306e-04 2.63014243e-03 -4.49641487e-04]
[0.00000000e+00 0.00000000e+00 0.00000000e+00
...,
0.00000000e+00 0.00000000e+00 0.00000000e+00]]]]

Acknowledgment

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Bibliography

Alexander Akoto-Danso works as an Assistant Research Scientist at the Ghana Space and Technology Institute. He holds a Master's in Computational Nuclear Science and Engineering (2011) and holds a Bsc. Mathematics. (2006) His interest is in High-performance Computing, Parallel computing, Computational Engineering, Digital Signal Processing, Climate modelling, Industrial and Systems engineering. He is a member of the Ghana Nuclear Society.

Felix Tetteh Madjitey works as an Assistant Research Scientist at the Ghana Space and Technology Institute. He holds a Master's in Computational Nuclear Science and Engineering and holds a Bsc. Physics. His research interest is in Structural and Mechanical Engineering.

Emmanuel Proven- Adzri is working as an Assistant Research Scientist at Ghana Space Science and Technology Institute. He has Masters in Computational Nuclear Science and Engineering and a BSc. in Physics. His research interest includes Instrumentation in Radio Astronomy, Methanol Masers, Radio Interferometry and Computational Modeling.

Theophilus Ansah-Narh is working as an Assistant Research Scientist at Ghana Space Science and Technology Institute. He has Masters in Statistics and Computational Nuclear Science and Engineering and also, BSc. Mathematics and Computer Science. His research interest includes Radio Interferometry, Stochastic Modeling, Bayesian Analysis, Computational Mathematics and Parallel Computing.

Marcellin Atemkeng holds a MSc Computer Sc. and BSc Mathematics and Computer Sc. His research interest includes Networking, Structure Documents, Workflow Systems, Simulation, Signal Correlation Algorithm Techniques, Synthesis Imaging in Radio Interferometry Arrays and Data Reduction.